



# The gravitational waves detector Advanced Virgo

Original slides courtesy of Julia Casanueva INFN Pisa

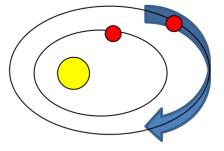
# Introduction

## Gravitational force

#### **Newton**, 1687 $\rightarrow$ Universal gravitation



$$F = G \frac{M \cdot m}{R^2}$$

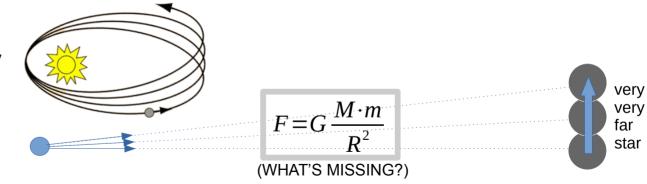


#### Pros

- Very nice correspondence with observations
- One theory for different phenomena
- Dramatic leap of culture and forecasting skills
- Please, stop me

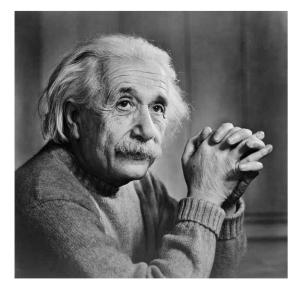
#### Cons

- Orbit of Mercury
- Try to guess



### Gravitational force

#### **Einstein**, 1915 $\rightarrow$ General Relativity

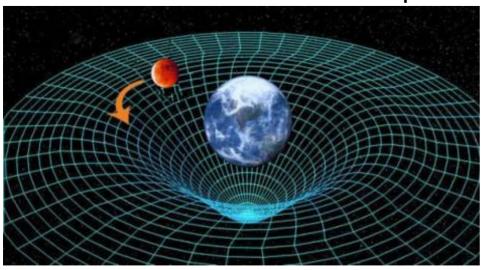


#### Some assumptions

- All repetitions of the same experiment in any free falling lab have the same results
- Time and space are deeply involved

#### Some consequences

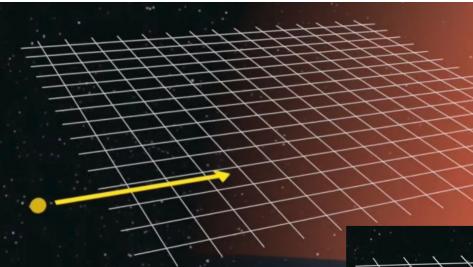
- Gravitation is not an interaction, but a property of space-time
  - Mass is the cause (or symptom) of a space-time distortion



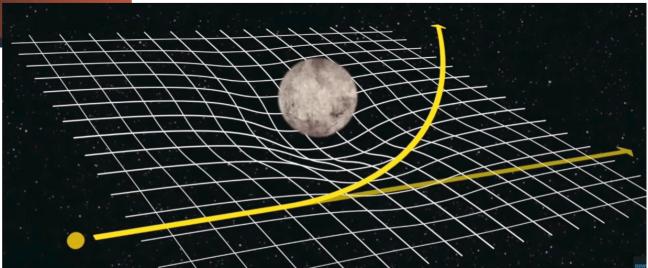
– Space-time fabric

### **General Relativity**

The presence of objects with mass modifies the space-time curvature

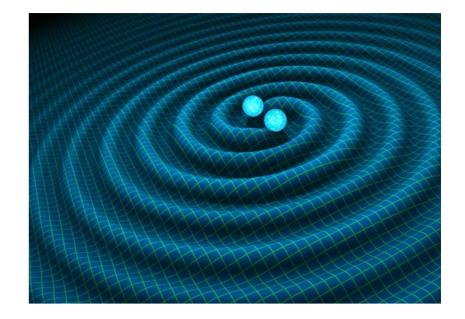


This curvature modifies the trajectory of other objects (and light)



# Gravitational waves

- Predicted by General Relativity
- Oscillations of the space-time fabric



- Produced by big masses accelerated
  - Emitted power: <u>asymmetric</u> mass distribution, <u>compact</u> and <u>relativistic.</u>
  - Distance: amplitude decreases as <u>1/r</u>

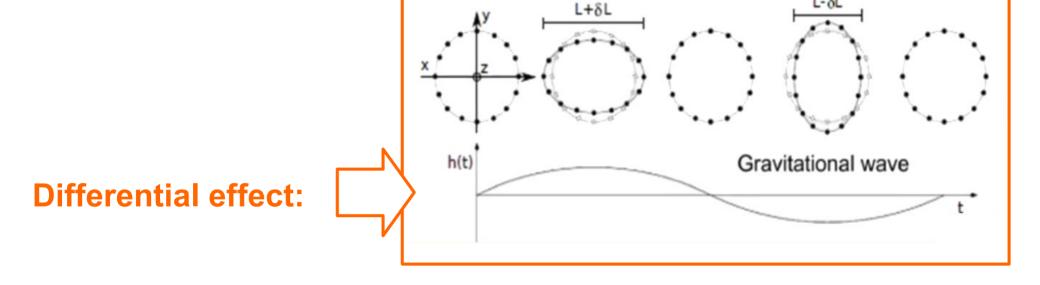
#### **Black Holes**



300.000 billions of billions of elephants in a car!

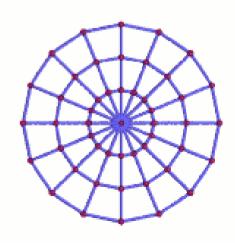
# Characteristics

- They travel at the speed of light
- They traverse matter unchanged
- How do we notice them?
  - They distort the space-time fabric around us  $\rightarrow$  for masses in *free fall*



Animation by Markus Pössel, www.einstein-online.info

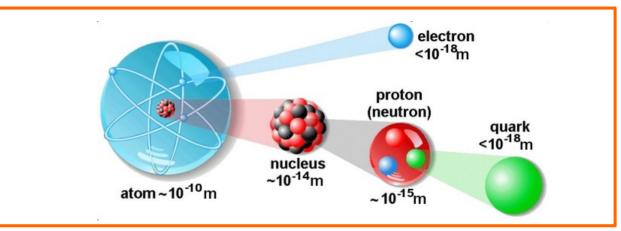
+ polarization



# Intensity of expected waves

- GW are local contractions/expansions of space.  $h = \frac{\text{variation of length}}{\text{length}}$  is a measure of their amplitude
- Far astronomic events  $\rightarrow$ h = 0.0000000000000000001 = 10<sup>-21</sup> on Earth
- With a reference length of 3 km  $\rightarrow \delta L = h \cdot L = 3 \cdot \frac{10^{-18}}{10^{-18}}$  m

Some interesting lengths in nature:



# Allegory

- Surface of the Oceans (all summed up):
  - $S = 3.6 \cdot 10^{14} \text{ m}^2$

• Volume of a Coke:

 $V = 0.33 \cdot 10^{-3} \text{ m}^{3}$ 

Ocean level rise after Coke can pouring:

$$= V/S = 10^{-18} m$$

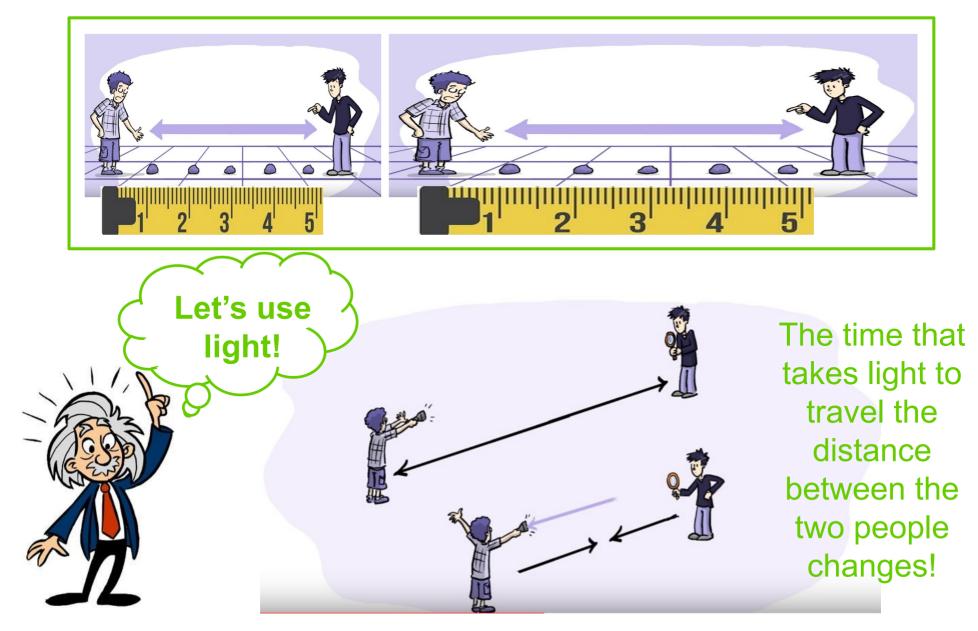
δL

Panic:
there are
two more
Cokes !

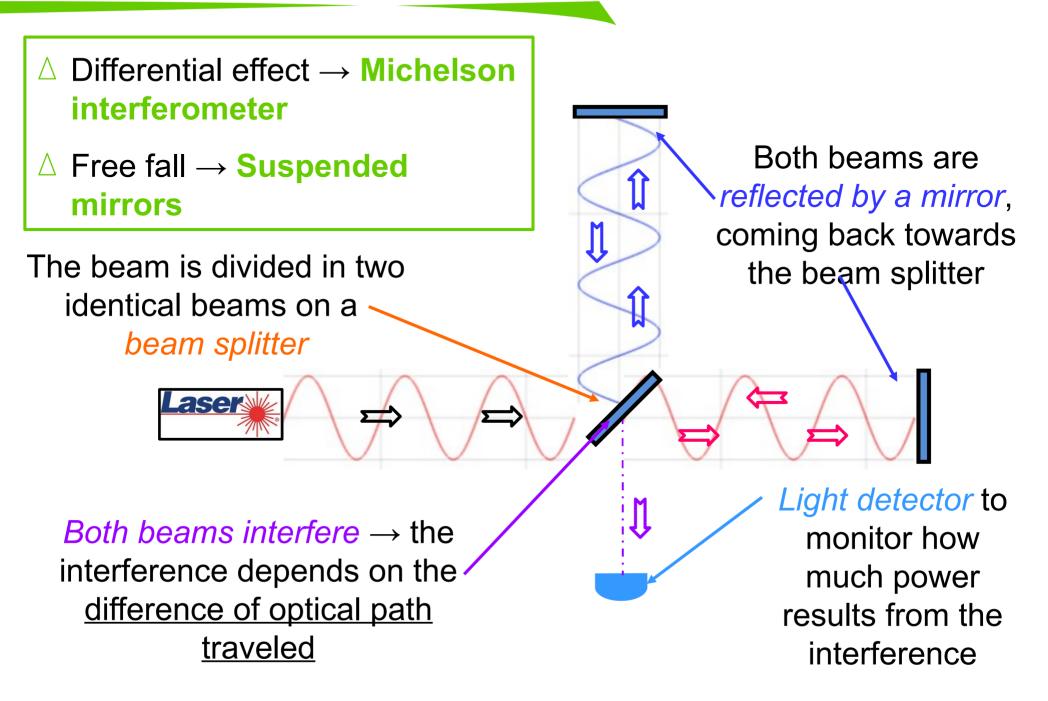
# **Detection principle**

#### How can we detect them?

 $\triangle$  We can't use a meter  $\rightarrow$  it will suffer as well the effects of the GW!

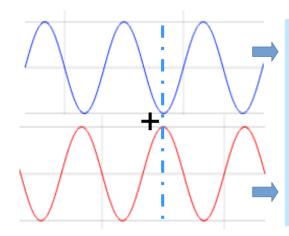


### Michelson interferometer

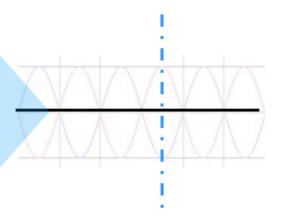


#### Interferences

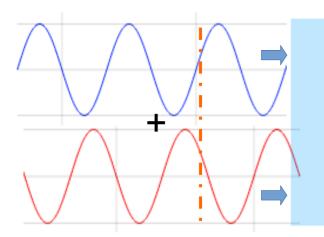
△ If both "arms" of the interferometer have the same length



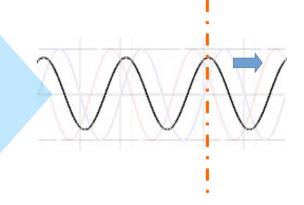
beams have opposed phase. Thus, when they overlap, they cancel in a perfect way → there is no light reaching the detector



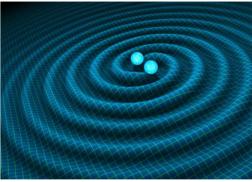
 ${\boldsymbol{\bigtriangleup}}$  If there is one arm longer than the other

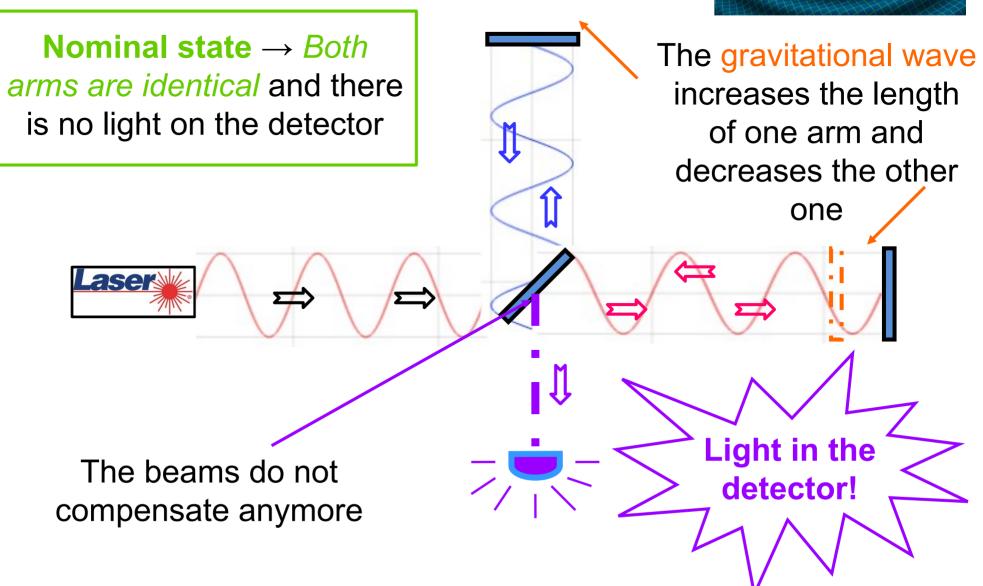


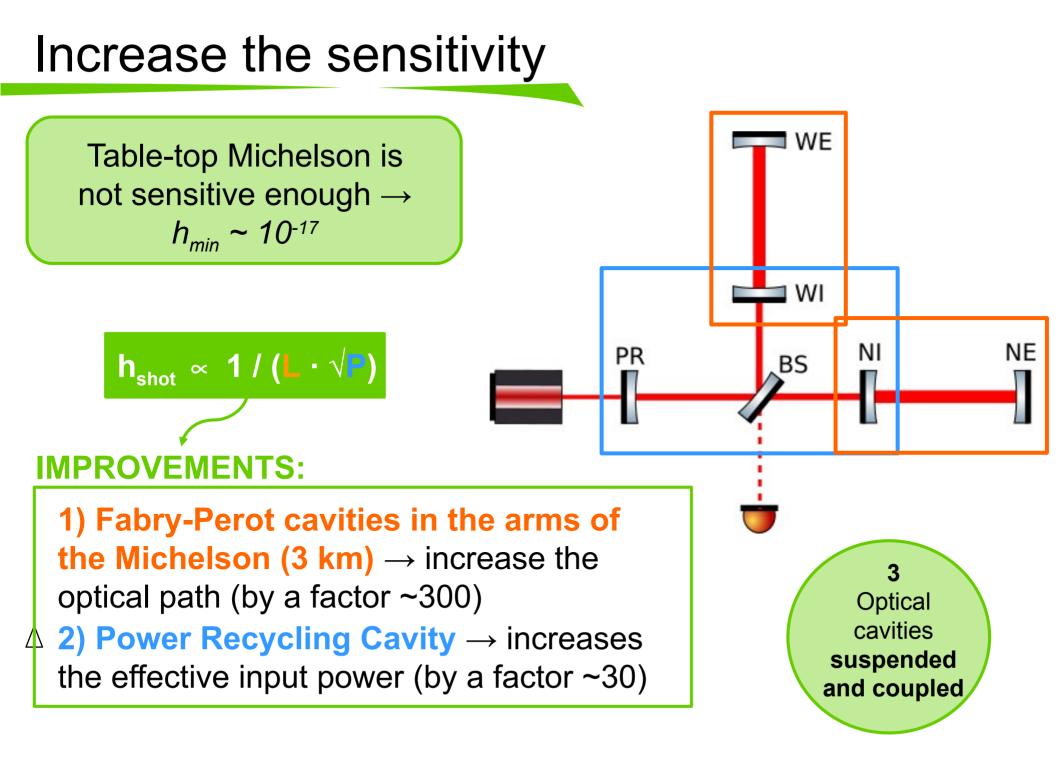
the sum of the two beams is different from zero → some light reaches the detector



## **Detection principle**





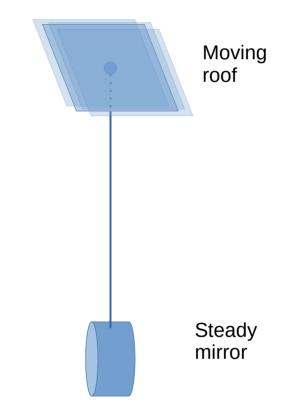


# Noise mitigation

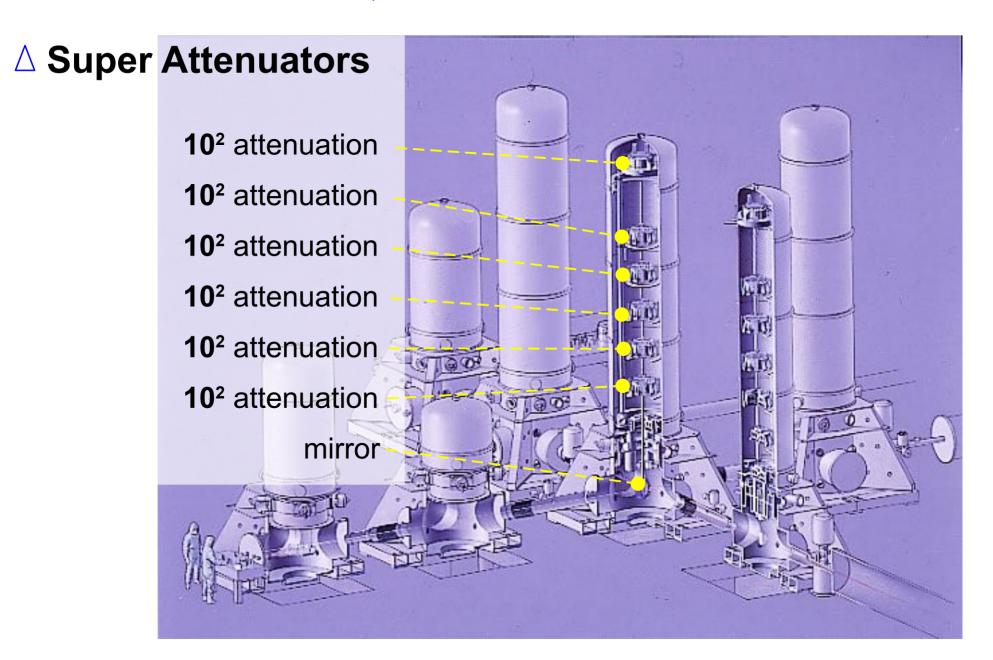
#### Noises

- △ Remember: the required accuracy is  $10^{-18}$  m out of 3 km → all fluctuations of earth, air, temperature, *etc*, are limiting
- △ An example: seismic vibrations (RMS amplitude, above 10 Hz): 4 · 10<sup>-8</sup> m

```
Mirrors have to be
more than 10<sup>10</sup> times steadier
than their hanging point
```



### Mechanical isolation of mirrors



# Limiting noises

△ Extreme techniques have been designed and implemented to *mitigate the different noise sources:* 

△ Seismic noise: at low frequencies the movement of the earth would be dominant  $\rightarrow$  **10**<sup>12</sup> of attenuation from 10 Hz

△ Pressure fluctuations: is necessary to work under vacuum → P =  $10^{-9}$  mbar (largest ultra-high-vacuum system in Europe)

△ Fluctuations of the laser: need a stable laser beam (v, alignment, P...) and "clean" (high content of TEM00)

 $\triangle$  Mirrors quality: low roughness  $\rightarrow$  10<sup>-10</sup> m RMS

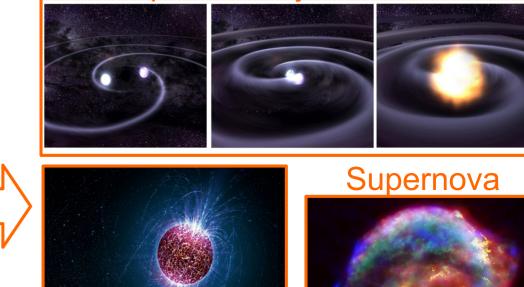
# Commissioning

### Sources of GW

#### WHY GWs ARE INTERESTING?

New source of astrophysical observations  $\rightarrow$  complementary to EM spectra + <u>new phenomena</u>

**Test General Relativity** 



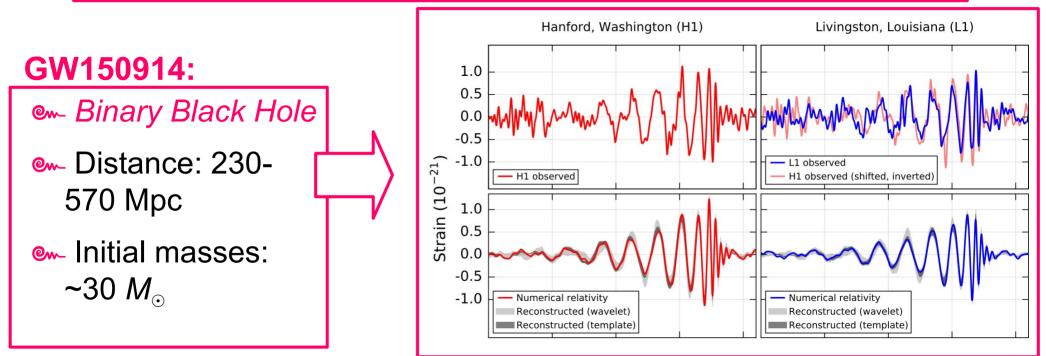
#### **Compact Binary Coalescence**

Some **sources** that emit gravitational waves strong enough to be detected by ground-based GW detectors

#### Rotating Neutron Star

## First detection: GW150914





#### GW151226:

en- Binary Black Holes

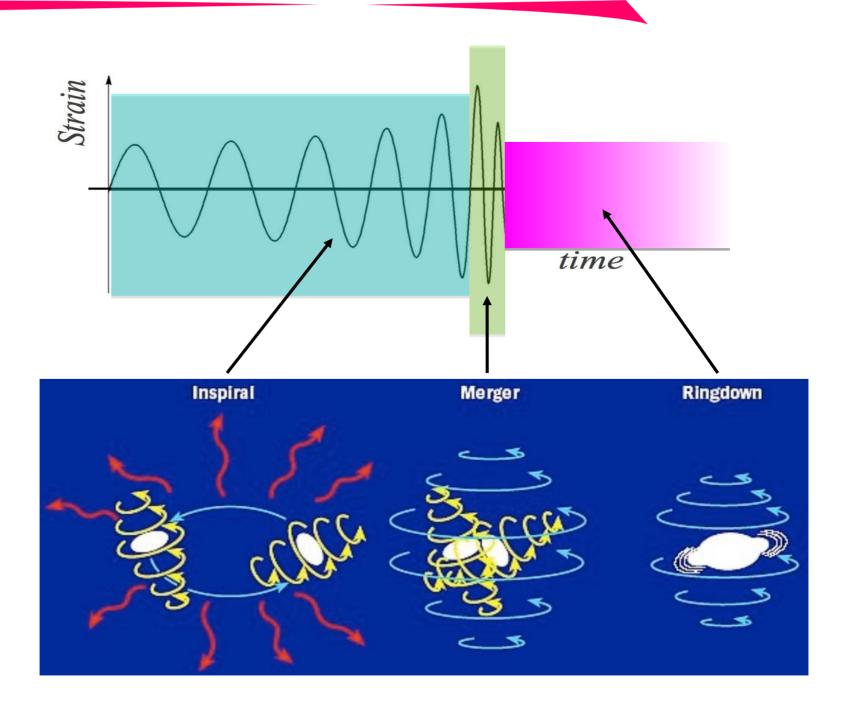
- Mpc Distance: 250-620 Mpc
- Initial masses: 11-23 and 5-10  $M_{\odot}$

#### GW170104:

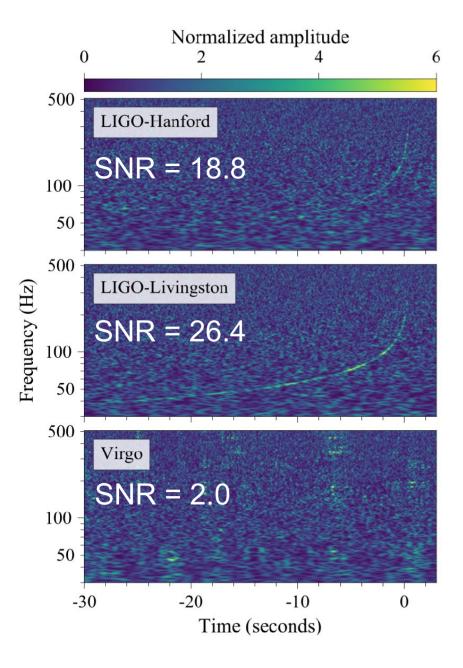
e Binary Black Holes

- Distance: 490-1318 Mpc
- Initial masses: 25-40 and 13-25  $M_{\odot}$

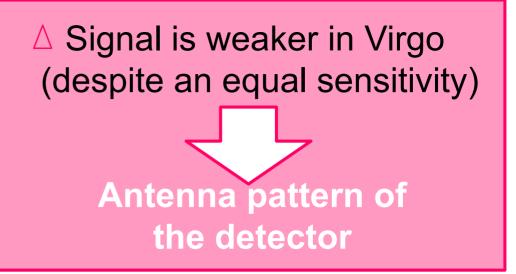
### Signal of a gravitational wave



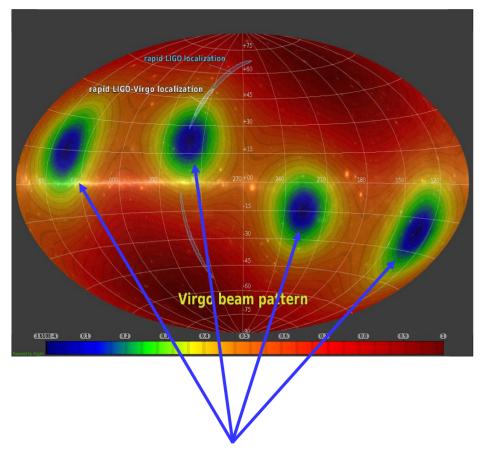
## First detection of a BNS: GW170817



- en- Binary Neutron Star
- Distance: 26-48 Mpc
- Initial masses: 1.36-2.26 and 0.86-1.36 M<sub>o</sub>
- Most energetic event  $(3M_{\odot}c^2)$
- Longest signal (~100 s)



#### Antenna pattern of the detector

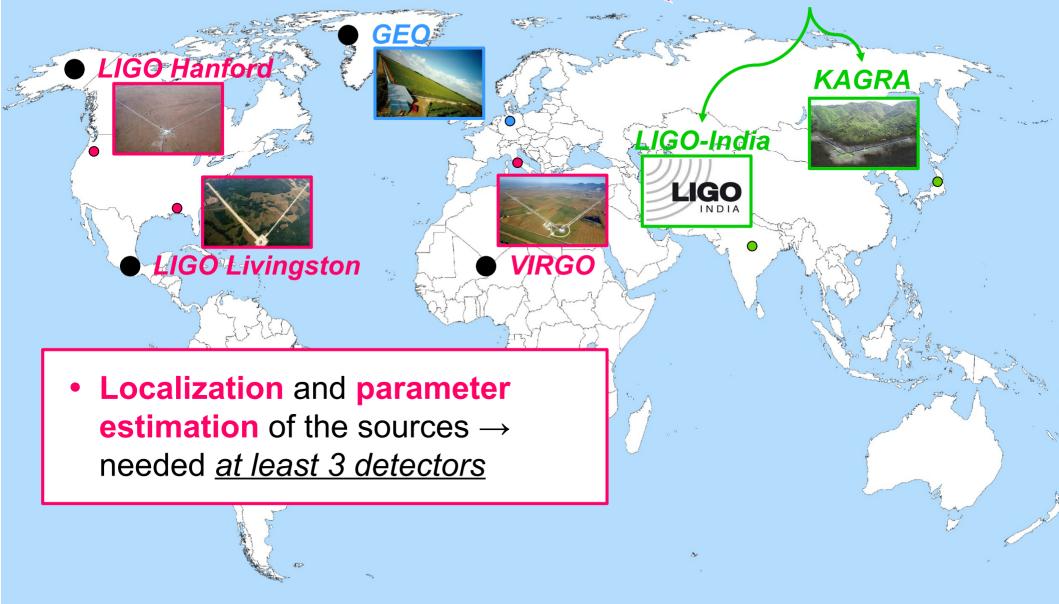


The detector is not equally sensitive in all directions

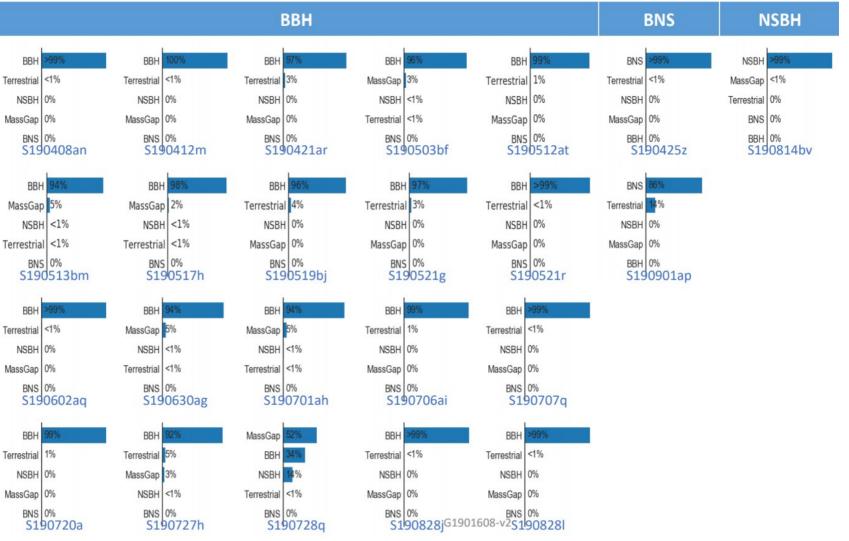
Sky localization only with LIGO (190 deg<sup>2</sup>) Sky localization with LIGO+Vįrgo (28 deg<sup>2</sup>) 5° 30 0° 15H 12h 9h 18h -30° -30°

#### GW detectors network

#### Future detectors



#### O3 observation period (last one) (Apr 2019 - Mar 2020)



#### More than 1 detection / week GW astronomy runs

# To be continued...